Scale invariance, T-duality and modified IIB supergravity Integrable eta-deformations



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Outline

- Deforming string sigma model
- Bosonic model
- Inclusion of fermions RR fields
- Scale invariance NSNS fields
- Modified type II equations
- Scale invariance RR fields
- Conclusions and future problems



Undeformed theory: symmetries and Lagrangian

Coset $PSU(2,2|4)/SO(4,1) \times SO(5)$





 $SU(2,2)/SO(4,1) \sim AdS_5$

 ${\rm SU}(4)/{\rm SO}(5)\sim {\rm S}^5$

coset element $\mathfrak{g}(\tau, \sigma)$ Current $A_{\alpha} \equiv -\mathfrak{g}^{-1}\partial_{\alpha}\mathfrak{g} \in \mathfrak{su}(2, 2|4)$ Realisation of $\mathfrak{su}(2, 2|4)$ in terms of 8×8 matrices

Undeformed theory: symmetries and Lagrangian

Z₄-grading:
$$f = \mathfrak{su}(2, 2|4) = f^{(0)} + f^{(1)} + f^{(2)} + f^{(3)}$$

$$\Omega(f^{(k)}) = i^k f^{(k)}$$
Projectors $P^{(k)}$, $d \equiv P^{(1)} + 2P^{(2)} - P^{(3)}$

$$\mathcal{L} = -\frac{g}{4} (\gamma^{\alpha\beta} - \epsilon^{\alpha\beta}) \operatorname{str} \left(A_{\alpha} \cdot d(A_{\beta}) \right)$$

This expression reproduces the expected Green – Schwarz Lagrangian for IIB superstring in $AdS_5 \times S^5$

Metsaev, Tseytlin '98

Deforming the theory

$$\mathcal{L}_{\eta} = -\frac{g}{4} (1 + \eta^{2}) (\gamma^{\alpha\beta} - \epsilon^{\alpha\beta}) \operatorname{str} \left(A_{\alpha} d \cdot \mathcal{O}^{-1} (A_{\beta}) \right)$$
$$d \equiv P^{(1)} + \frac{2}{1 - \eta^{2}} P^{(2)} - P^{(3)}$$
$$\mathcal{O} \equiv 1 - \eta R_{\mathfrak{g}} \circ d$$
$$R_{\mathfrak{g}}(\cdot) \equiv \operatorname{Adj}_{\mathfrak{g}^{-1}} \circ R \circ \operatorname{Adj}_{\mathfrak{g}}(\cdot) = \mathfrak{g}^{-1} R(\mathfrak{g}(\cdot)\mathfrak{g}^{-1})\mathfrak{g}$$

modified classical Yang-Baxter equation (mcYBe)

[RM, RN] - R([RM, N] + [M, RN]) = [M, N]

Delduc, Magro, Vicedo '13

 $\eta \in [0,1[$

Generic properties if the deformed models

Classically integrable

 \mathbf{M} -symmetry

 \checkmark Hidden symmetry $\mathrm{PSU}_q(2,2|4)$

Material R's, corresponding to Dynkin diagrams

For the rest of the talk we choose R which corresponds to the standard Dynkin diagram



$$R(M_{jk}) = \begin{cases} -i M_{jk}, & j < k \\ 0, & j = k \\ +i M_{jk}, & j > k \end{cases}$$

Other interesting choices are possible!

Bosonic sigma-model

$$S = -\frac{\tilde{g}}{2} \int d\tau d\sigma \left(\gamma^{\alpha\beta} \partial_{\alpha} X^{M} \partial_{\beta} X^{N} G_{MN} - \epsilon^{\alpha\beta} \partial_{\alpha} X^{M} \partial_{\beta} X^{N} B_{MN} \right) \qquad \qquad \varkappa = \frac{2\eta}{1 - \eta^{2}}$$

$$\tilde{g} \equiv g \sqrt{1 + \varkappa^{2}}$$

$$f_{\pm}(x) = 1 \pm x^{2}$$

$$ds^{2} = -\frac{f_{+}(\rho)}{f_{-}(\varkappa\rho)}dt^{2} + \frac{1}{f_{+}(\rho)f_{-}(\varkappa\rho)}d\rho^{2} + \rho^{2}d\Theta_{3}^{\rho} + \frac{f_{-}(r)}{f_{+}(\varkappa r)}d\phi^{2} + \frac{1}{f_{-}(r)f_{+}(\varkappa r)}dr^{2} + r^{2}d\Theta_{3}^{r}$$

$$\frac{1}{1+\varkappa^2 q^4 \sin^2 \zeta} (d\zeta^2 + \cos^2 \zeta d\psi_1^2) + \sin^2 \zeta d\psi_2^2$$

$$d\Theta_{3}^{\rho} \equiv \frac{1}{1 + \varkappa^{2} \rho^{4} \sin^{2} \zeta} (d\zeta^{2} + \cos^{2} \zeta d\psi_{1}^{2}) + \sin^{2} \zeta d\psi_{2}^{2}$$
$$d\Theta_{3}^{r} \equiv \frac{1}{1 + \varkappa^{2} r^{4} \sin^{2} \xi} (d\xi^{2} + \cos^{2} \xi d\chi_{1}^{2}) + \sin^{2} \xi d\chi_{2}^{2}$$

$$B = \varkappa \Big(\frac{\rho^4 \sin 2\zeta}{1 + \varkappa^2 \rho^4 \sin^2 \zeta} d\psi_1 \wedge d\zeta - \frac{r^4 \sin 2\xi}{1 + \varkappa^2 r^4 \sin^2 \xi} d\chi_1 \wedge d\xi \Big)$$

 $\stackrel{\scriptscriptstyle \odot}{\scriptstyle \backsim} {\rm The \ limit} \ \ \varkappa \to 0 \Longrightarrow AdS_5 \times S^5$

arphi The metric and B-field are invariant under shifts of $t,\phi,\psi_1,\psi_2,\chi_1,\chi_2$

$$\mathbf{\hat{e}}$$
 The ranges $0\leq
ho\leqrac{1}{arkappa}\,,\,\,\,\,\,0\leq r\leq 1$

§ The deformed AdS is singular at ~
ho=1/arkappa

Borsato, Frolov and G.A. '13















RR fields

Green-Schwarz Lagrangian at quadratic order in fermions

 $\mathcal{L}_{\theta^2} = -\frac{g}{2} \left(\gamma^{\alpha\beta} \delta^{IJ} + \epsilon^{\alpha\beta} \sigma_3^{IJ} \right) i \bar{\theta}_I e_{\alpha}^m \gamma_m D_{\beta}^{JK} \theta_K$



Extract $e^{\Phi}F^{(n)}$, plug in SUGRA equations for $F^{(n)} \implies$ lst order differential equation for Φ !

Green-Schwarz Lagrangian at quadratic order in fermions

Reaching the canonical form of the GS Lagrangian

 $\theta \to U\theta, \qquad X^M \to X^M + \bar{\theta}f^M\theta$ $U, f^M : 32 \times 32$ matrices, functions of X

$$S_{\text{bos}}(X) \to S_{\text{bos}}(X) + \delta S_{\text{bos}}(X, \overline{\theta}f\theta)$$

 $\bar{\theta}_I M_{IJ} \theta_J \to \bar{\theta}_I \ \left(\bar{U} M U \right)_{IJ} \theta_J$

Guiding principle: find U and f^M to get a canonical term with ∂

$$-\frac{\tilde{g}}{2}\left(\gamma^{\alpha\beta}\delta^{IJ}+\epsilon^{\alpha\beta}\sigma_{3}^{IJ}\right) \ i\bar{\theta}_{I} \ \tilde{e}_{\alpha}^{m}\gamma_{m}\partial_{\beta}\theta_{J}$$

 \widetilde{e}_{M}^{m} is the deformed vielbein

We found such a field redefinition! The terms with ω and $H_{\mu\nu\rho}$ came out correctly!

Results for the RR-sector (ABF background)

$$e^{\phi}F_1 = -4\varkappa^2 \sqrt{1+\varkappa^2} \ c_F^{-1} \ \rho^3 \sin\zeta, \qquad e^{\phi}F_6 = -4\varkappa^2 \sqrt{1+\varkappa^2} \ c_F^{-1} \ r^3 \sin\xi,$$

$$e^{\phi}F_{014} = +4\varkappa\sqrt{1+\varkappa^2} \ c_F^{-1} \ \rho^2 \sin\zeta,$$

$$e^{\phi}F_{569} = -4\varkappa\sqrt{1+\varkappa^2} \ c_F^{-1} \ r^2 \sin\xi,$$

$$e^{\phi}F_{046} = -4\varkappa^3\sqrt{1+\varkappa^2} \ c_F^{-1} \ \rho r^3 \sin\xi,$$

$$e^{\phi}F_{159} = -4\varkappa^3\sqrt{1+\varkappa^2} \ c_F^{-1} \ \rho^3 r \sin\zeta,$$

$$e^{\phi}F_{123} = -4\varkappa\sqrt{1+\varkappa^2} c_F^{-1} \rho,$$

$$e^{\phi}F_{678} = +4\varkappa\sqrt{1+\varkappa^2} c_F^{-1} r,$$

$$e^{\phi}F_{236} = +4\varkappa^3\sqrt{1+\varkappa^2} c_F^{-1} \rho^2 r^3 \sin\zeta\sin\xi,$$

$$e^{\phi}F_{178} = -4\varkappa^3\sqrt{1+\varkappa^2} c_F^{-1} \rho^3 r^2 \sin\zeta\sin\xi,$$

$$e^{\phi}F_{01234} = +4\sqrt{1+\varkappa^2} c_F^{-1},$$

$$e^{\phi}F_{01459} = +4\varkappa^2\sqrt{1+\varkappa^2} c_F^{-1}\rho^2 r\sin\zeta,$$

$$e^{\phi}F_{04569} = -4\varkappa^2\sqrt{1+\varkappa^2} c_F^{-1}\rho r^2\sin\xi,$$

$$e^{\phi}F_{02346} = +4\varkappa^{4}\sqrt{1+\varkappa^{2}} c_{F}^{-1}\rho^{3}r^{3}\sin\zeta\sin\xi,$$

$$e^{\phi}F_{01478} = +4\varkappa^{2}\sqrt{1+\varkappa^{2}} c_{F}^{-1}\rho^{2}r^{2}\sin\zeta\sin\xi,$$

$$e^{\phi}F_{04678} = +4\varkappa^{2}\sqrt{1+\varkappa^{2}} c_{F}^{-1}\rho r.$$

Х

$$c_F = \sqrt{1 - \varkappa^2 \rho^2} \sqrt{1 + \varkappa^2 \rho^4 \sin^2 \zeta} \sqrt{1 + \varkappa^2 r^2} \sqrt{1 + \varkappa^2 r^4 \sin^2 \xi}$$

SUGRA eoms are not satisfied!

Maldacena-Russo background is reproduced!

Maldacena-Russo background

$$\begin{split} &\varkappa \to 0 \,, \quad \rho \to \rho/\sqrt{\varkappa} \,, \quad t \to \sqrt{\varkappa}t \,, \quad \zeta \to \zeta_0 + \sqrt{\varkappa}\zeta \\ &\psi_1 \to \sqrt{\varkappa} \,\psi_1/\cos\zeta_0 \,, \quad \psi_2 \to \sqrt{\varkappa} \,\psi_2/\sin\zeta_0 \end{split}$$

$$ds^{2} = \rho^{2}(-dt^{2} + d\psi_{2}^{2}) + \frac{\rho^{2}(d\zeta^{2} + d\psi_{1}^{2})}{1 + \rho^{4}\sin^{2}\zeta_{0}} + \frac{d\rho^{2}}{\rho^{2}},$$
$$B_{a} = \frac{\rho^{4}\sin\zeta_{0}}{1 + \rho^{4}\sin^{2}\zeta_{0}}d\psi_{1} \wedge d\zeta$$

[Matsumoto, Yoshida '14] [Frolov, unpublished '14]

$$\log c_F^{-1} \implies \qquad \phi = \phi_0 - \frac{1}{2} \log(1 + \rho^4 \sin^2 \zeta_0)$$

$$F_{014} \implies \qquad F_3 = 4\rho^3 \sin \zeta_0 e^{-\phi_0} dt \wedge d\psi_2 \wedge d\rho$$

 $F_{01234} \implies F_{5} = 2e^{-\phi_{0}} \left(\frac{2\rho^{3}}{1 + \rho^{4} \sin^{2} \zeta_{0}} dt \wedge d\psi_{2} \wedge d\psi_{1} \wedge d\zeta \wedge d\rho + r^{3} \sin 2\xi d\phi \wedge d\phi_{2} \wedge d\phi_{1} \wedge d\xi \wedge dr \right)$

[Maldacena, Russo '99]

Kappa-symmetry

The deformed model has kappa-symmetry

32-dim spinors $K_{\alpha I}$ to parameterise kappa-symmetry transf. $\Theta_I = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \otimes \theta_I$ and Γ_m are 32 × 32 Gamma matrices

The same field redefinitions bring the kappa-symmetry transformations to the standard form:

$$\begin{split} \delta_{\kappa} X^{M} &= -\frac{i}{2} \; \bar{\Theta}_{I} \delta^{IJ} \Gamma^{M} \delta_{\kappa} \Theta_{J} + \mathcal{O}(\Theta^{3}), \\ \delta_{\kappa} \Theta_{I} &= -\frac{1}{4} (\delta^{IJ} \gamma^{\alpha\beta} - \sigma_{3}^{IJ} \epsilon^{\alpha\beta}) \Gamma_{\beta} K_{\alpha J} + \mathcal{O}(\Theta^{2}), \end{split}$$

RR-couplings can be read off from the variation of the world-sheet metric and they are the same as found before!

$$\frac{\delta_{\kappa}\gamma^{\alpha\beta}}{\Gamma^{IJ\,\alpha\alpha'}} = 2i \, \Pi^{IJ\,\alpha\alpha'} \Pi^{JK\,\beta\beta'} \, \bar{K}_{I\alpha'} D^{KL}_{\beta'} \Theta_L + \mathcal{O}(\Theta^{4})$$
$$\Pi^{IJ\,\alpha\alpha'} \equiv \frac{\delta^{IJ}\gamma^{\alpha\alpha'} + \sigma^{IJ}_3 \epsilon^{\alpha\alpha'}}{2}$$

Contradiction with [Bergshoeff, Sezgin, Townsend, '86] and [Grisaru, Howe, Mezincescu, Nilsson, Townsend '86]?

kappa-symmetry implies supergravity constraints (eoms?)

Differential constraints can start showing up only from variations of quartic action in fermions!

HT background in the T-dualized model

$$\hat{ds}^{2} = -\frac{1-\varkappa^{2}\rho^{2}}{1+\rho^{2}}d\hat{t}^{2} + \frac{d\rho^{2}}{(1+\rho^{2})(1-\varkappa^{2}\rho^{2})} + \frac{d\hat{\psi}_{1}^{2}}{\rho^{2}\cos^{2}\zeta} + (\rho d\zeta + \varkappa\rho\tan\zeta\,d\hat{\psi}_{1})^{2} + \frac{d\hat{\psi}_{2}^{2}}{\rho^{2}\sin^{2}\zeta} + \frac{1+\varkappa^{2}r^{2}}{(1-r^{2})(1+\varkappa^{2}r^{2})} + \frac{d\hat{\phi}_{1}^{2}}{r^{2}\cos^{2}\xi} + (rd\xi - \varkappa r\tan\xi d\hat{\phi}_{1})^{2} + \frac{d\hat{\phi}_{2}^{2}}{r^{2}\sin^{2}\xi}$$

 $\hat{B} = 0, \qquad \hat{\mathcal{F}}_1 = 0 = \hat{\mathcal{F}}_3$

$$\hat{\phi} = \phi_0 - 4\varkappa(\hat{t} + \hat{\varphi}) - 2\varkappa(\hat{\psi}_1 - \hat{\phi}_1) + \log\frac{(1 - \varkappa^2 \rho^2)^2 (1 + \varkappa^2 r^2)^2}{\rho^2 r^2 \sqrt{1 + \rho^2} \sqrt{1 - r^2} \sin 2\zeta \sin 2\xi}$$



- An exact solution of IIB sugra (metric is non-diagonal)
- $\hat{\mathbf{F}}$ Flux $\hat{\mathcal{F}}_5$ is imaginary
- Dilaton is linear in isometric coordinates
- Formally T-dualizing back is not possible!

Scale invariance of the eta-deformed model NSNS-sector

Scale vs Weyl Invariance for the bosonic sigma model

Scale invariance: δq

 $\delta g_{\mu\nu} = \epsilon g_{\mu\nu}$

$$\int T^{\mu}_{\mu} = 0$$
 or locally $T^{\mu}_{\mu} = \partial_{\mu}K^{\mu}$

Weyl invariance:

 $\delta g_{\mu\nu} = \epsilon(x) g_{\mu\nu}$

 $T^{\mu}_{\mu} = 0$

Scale invariance conditions for the bosonic sigma model:

$$\beta_{mn}^G \equiv R_{mn} - \frac{1}{4} H_{mkl} H_n^{\ kl} = -D_m X_n - D_n X_m$$
$$\beta_{mn}^B \equiv \frac{1}{2} D^k H_{kmn} + \partial_m Y_n - \partial_n Y_m$$

Weyl invariance conditions for the bosonic sigma model:

$$X_m = \partial_m \phi$$
, $Y_m = 0$ ϕ is a dilation

$$\partial_m \beta^\phi = 0 , \qquad \beta^\phi \equiv R - \frac{1}{12} H_{mnk}^2 + 4D^2 \phi - 4\partial^m \phi \partial_m \phi$$

Central charge identity (Curci-Paffuti identity)

 $T = \beta_{mn}^G \partial_a x^m \partial^a x^n + \beta_{mn}^B \epsilon^{ab} \partial_a x^m \partial_b x^n, \qquad T = \nabla^a N_a, \qquad N_a = 2(X_m \partial_a x^m + \epsilon_a^{\ b} Y_m \partial_b x^m)$ This cannot be cancelled by a local counterterm unless $X_m = \partial_m \phi, \ Y_m = 0$

Scale invariance conditions for the GS superstring

$$\mathcal{T}_{mn} \equiv \frac{1}{2} \mathcal{F}_m \mathcal{F}_n + \frac{1}{4} \mathcal{F}_{mpq} \mathcal{F}_n^{pq} + \frac{1}{4 \times 4!} \mathcal{F}_{mpqrs} \mathcal{F}_n^{pqrs} - \frac{1}{2} \mathcal{G}_{mn} \left(\frac{1}{2} \mathcal{F}_k \mathcal{F}^k + \frac{1}{12} \mathcal{F}_{kpq} \mathcal{F}^{kpq} \right)$$

$$\beta_{mn}^G \equiv R_{mn} - \frac{1}{4} H_{mkl} H_n^{\ kl} - \mathcal{T}_{mn} = -D_m X_n - D_n X_m$$

$$\beta_{mn}^{B} \equiv \frac{1}{2} D^{k} H_{kmn} + \mathcal{K}_{mn} = X^{k} H_{kmn} + \partial_{m} Y_{n} - \partial_{n} Y_{m}$$
$$\mathcal{K}_{mn} \equiv \frac{1}{2} \mathcal{F}^{k} \mathcal{F}_{kmn} + \frac{1}{12} \mathcal{F}_{mnklp} \mathcal{F}^{klp}$$

$$\mathcal{F}_m \equiv e^{\phi} F_m , \quad \mathcal{F}_{mnk} \equiv e^{\phi} F_{mnk} , \quad \mathcal{F}_{mnklp} \equiv e^{\phi} F_{mnklp}$$

1)

For $X_m = \partial_m \phi$, $Y_m = 0$ these equations follow from IIB supergravity action

2) For RR-fields UV finiteness of the 2d sigma model should imply 2nd order differential equations:

$$\beta_{k_1\dots k_s}^{\mathcal{F}} \equiv \frac{1}{2} D^2 \mathcal{F}_{k_1\dots k_s} + \dots = X^m \partial_m \mathcal{F}_{k_1\dots k_s} + \sum_i \mathcal{F}_{k_1\dots m\dots k_s} \partial_{k_i} X^m$$

X & Y of the eta-deformed model

$$\begin{split} X &\equiv X_m dx^m = c_0 \frac{1+\rho^2}{1-\varkappa^2 \rho^2} dt + c_1 \rho^2 \sin^2 \zeta \, d\psi_2 + c_2 \frac{\rho^2 \cos^2 \zeta}{1+\varkappa^2 \rho^4 \sin^2 \zeta} d\psi_1 \\ &+ c_3 \frac{1-r^2}{1+\varkappa^2 r^2} d\varphi + c_4 r^2 \sin^2 \xi \, d\phi_2 + c_5 \frac{r^2 \cos^2 \xi}{1+\varkappa^2 r^4 \sin^2 \xi} d\phi_1 \\ &+ \frac{\varkappa^2 \rho^4 \sin 2\zeta}{2(1+\varkappa^2 \rho^4 \sin^2 \zeta)} d\zeta + \frac{1}{\rho} (1 - \frac{3}{1-\varkappa^2 \rho^2} + \frac{2}{1+\varkappa^2 \rho^4 \sin^2 \zeta}) d\rho \\ &+ \frac{\varkappa^2 r^4 \sin 2\xi}{2(1+\varkappa^2 r^4 \sin^2 \xi)} d\xi + \frac{1}{r} (1 - \frac{3}{1+\varkappa^2 r^2} + \frac{2}{1+\varkappa^2 r^4 \sin^2 \xi}) dr \end{split}$$

Isometry

$$X_m = I_m + Z_m$$
, $D_m I_n + D_n I_m = 0$, $D^m I_m = 0$

$$Y \equiv Y_m dx^m = 4\varkappa \frac{1+\rho^2}{1-\varkappa^2 \rho^2} dt + 2\varkappa \frac{\rho^2 \cos^2 \zeta}{1+\varkappa^2 \rho^4 \sin^2 \zeta} d\psi_1 + 4\varkappa \frac{1-r^2}{1+\varkappa^2 r^2} d\varphi - 2\varkappa \frac{r^2 \cos^2 \xi}{1+\varkappa^2 r^4 \sin^2 \xi} d\phi_1 + \frac{\varkappa^2 \rho^4 \sin 2\zeta}{2(1+\varkappa^2 \rho^4 \sin^2 \zeta)} d\zeta + \frac{1}{\rho} \left(1 - \frac{3}{1-\varkappa^2 \rho^2} + \frac{2(\varkappa^{-1} c_2 - 1)}{1+\varkappa^2 \rho^4 \sin^2 \zeta}\right) d\rho + \frac{\varkappa^2 r^4 \sin 2\xi}{2(1+\varkappa^2 r^4 \sin^2 \xi)} d\xi + \frac{1}{r} \left(1 - \frac{3}{1+\varkappa^2 r^2} - \frac{2(\varkappa^{-1} c_5 + 1)}{1+\varkappa^2 r^4 \sin^2 \xi}\right) dr$$

For

$$c_0 = c_3 = 4\varkappa$$
, $c_1 = c_4 = 0$, $c_2 = -c_5 = 2\varkappa$
 $Y_m = X_m$

Surprisingly, for these coefficients
$$\bar{\beta}^X \equiv R - \frac{1}{12}H_{mnk}^2 + 4D_kX^k - 4X_kX^k = 0$$
 $\partial_m\bar{\beta}^X = 0$

ABF analog of the dilation



Solution:

$$X_m = Y_m = I_m + Z_m = \partial_m \phi + (G_{km} + B_{km})I^k$$

$$\phi = \frac{1}{2} \log \frac{(1 - \kappa^2 \rho^2)^3 (1 + \kappa^2 r^2)^3}{(1 + \kappa^2 \rho^4 \sin^2 \zeta)(1 + \kappa^2 r^4 \sin^2 \xi)}$$
Precisely the isometric part of the HT dilation $\hat{\phi}$ under the standard T-duality

Modified type II equations RR-sector

[Frolov, Hoare, Roiban Tseytlin and G.A. '15]

Standard IIB sugra equations

Introduce
$$Z = Z_m dx^m = d\phi$$

 $\mathcal{F} = e^{\phi} F$

1)
$$D^{m}\mathcal{F}_{m} - Z^{m}\mathcal{F}_{m} - \frac{1}{6}H^{mnp}\mathcal{F}_{mnp} = 0, \qquad d\mathcal{F}_{1} - Z \wedge \mathcal{F}_{1} = 0,$$

2)
$$D^{p}\mathcal{F}_{pmn} - Z^{p}\mathcal{F}_{pmn} - \frac{1}{6}H^{pqr}\mathcal{F}_{mnpqr} = 0, \qquad d\mathcal{F}_{3} - Z \wedge \mathcal{F}_{3} + H_{3} \wedge \mathcal{F}_{1} = 0,$$

3)
$$D^{r}\mathcal{F}_{rmnpq} - Z^{r}\mathcal{F}_{mnpq} + \frac{1}{36}\varepsilon_{mnpqrstuvw}H^{rst}\mathcal{F}^{uvw} = 0, \quad d\mathcal{F}_{5} - Z \wedge \mathcal{F}_{5} + H_{3} \wedge \mathcal{F}_{3} = 0.$$

Dynamical equations Bianchi identities

$$\star \mathcal{F}_5 = \mathcal{F}_5$$

Self-duality of the five- form

I-modified sugra equations

1)

$$D^{m}\mathcal{F}_{m} - Z^{m}\mathcal{F}_{m} - \frac{1}{6}H^{mnp}\mathcal{F}_{mnp} = 0 , \qquad I^{m}\mathcal{F}_{m} = 0 ,$$
$$(d\mathcal{F}_{1} - Z \wedge \mathcal{F}_{1})_{mn} - I^{p}\mathcal{F}_{mnp} = 0 .$$

2)

$$D^{p}\mathcal{F}_{pmn} - Z^{p}\mathcal{F}_{pmn} - \frac{1}{6}H^{pqr}\mathcal{F}_{mnpqr} - (I \wedge \mathcal{F}_{1})_{mn} = 0 ,$$

$$(d\mathcal{F}_{3} - Z \wedge \mathcal{F}_{3} + H_{3} \wedge \mathcal{F}_{1})_{mnpq} - I^{r}\mathcal{F}_{mnpqr} = 0 .$$

3)
$$D^{r}\mathcal{F}_{rmnpq} - Z^{r}\mathcal{F}_{rmnpq} + \frac{1}{36}\varepsilon_{mnpqrstuvw}H^{rst}\mathcal{F}^{uvw} - (I \wedge \mathcal{F}_{3})_{mnpq} = 0 ,$$
$$(d\mathcal{F}_{5} - Z \wedge \mathcal{F}_{5} + H_{3} \wedge \mathcal{F}_{3})_{mnpqrs} + \frac{1}{6}\varepsilon_{mnpqrstuvw}I^{t}\mathcal{F}^{uvw} = 0 .$$

Surprisingly, there exists combinations of the above equations that depends on Z and I through the combination Z = X + I only:

$$D^{m}\mathcal{F}_{m} - X^{m}\mathcal{F}_{m} - \frac{1}{6}H^{mnp}\mathcal{F}_{mnp} = 0 ,$$

$$D^{p}\mathcal{F}_{pmn} - X^{p}\mathcal{F}_{pmn} - \frac{1}{6}H^{pqr}\mathcal{F}_{mnpqr} + (d\mathcal{F}_{1} - X \wedge \mathcal{F}_{1})_{mn} = 0 ,$$

$$D^{r}\mathcal{F}_{rmnpq} - X^{r}\mathcal{F}_{rmnpq} + \frac{1}{36}\varepsilon_{mnpqrstuvw}H^{rst}\mathcal{F}^{uvw} + (d\mathcal{F}_{3} - X \wedge \mathcal{F}_{3} + H_{3} \wedge \mathcal{F}_{1})_{mnpq} = 0$$

These equations are sufficient to derive 2nd order equations which should express scale invariance conditions

Equations for RR-fields as scale invariance conditions

Requirements:

- (i) vanish on the supergravity equations with $X = d\phi$, Y = 0
- (ii) depend on Z and I through X = Z + I
- (iii) depend on X through Lie derivatives

Compact form of the *I*-modified equations

$$d\mathcal{F}_{2n+1} - Z \wedge \mathcal{F}_{2n+1} + H_3 \wedge \mathcal{F}_{2n-1} - \star (I \wedge \star \mathcal{F}_{2n+3}) = 0 , \qquad n = -1, 0, \dots ,$$

$$d \star \mathcal{F}_{2n+1} - Z \wedge \star \mathcal{F}_{2n+1} - H_3 \wedge \star \mathcal{F}_{2n+3} + \star (I \wedge \mathcal{F}_{2n-1}) = 0 , \qquad n = 0, 1, \dots .$$

 $n = -1: \quad \star (I \wedge \star \mathcal{F}_1) = I^m \mathcal{F}_m = 0$

Act on the first equation with $\star d \star$ and on the second with $d \star$

Equations for RR fields as scale invariance conditions

$$\begin{array}{l} \textbf{1)} \qquad \frac{D^2 \mathcal{F}_m - R_{mn} \mathcal{F}^n + \frac{1}{4} (R - \frac{3}{4} H^2) \mathcal{F}_m}{+ \frac{1}{2} H^{pnk} H_{mpn} \mathcal{F}_k - \frac{1}{6} D_m H^{pnk} \mathcal{F}_{pnk} - \frac{1}{2} H^{pnk} D_p \mathcal{F}_{nkm}} \\ = 2(X^p D_p \mathcal{F}_m + D_m X^p \mathcal{F}_p) + \beta_{mn}^G \mathcal{F}^n - \frac{1}{2} \beta_{nk}^B \mathcal{F}^{nk}_m . \end{array}$$

$$\begin{array}{l} \textbf{2)} \qquad D^{2}\mathcal{F}_{nkm} - R_{a[n}\mathcal{F}^{a}_{\ \ km]} + R_{ab[nk}\mathcal{F}^{ab}_{\ \ m]} + \frac{1}{4}(R - \frac{3}{4}H^{2})\mathcal{F}_{nkm} \\ \\ \hline + \frac{1}{2}H^{abc}H_{ab[n}\mathcal{F}_{km]c} - \frac{1}{2}H^{abc}H_{a[nk}\mathcal{F}_{m]bc} \\ \\ + D^{a}H_{a[nk}\mathcal{F}_{m]} + H_{a[nk}D^{a}\mathcal{F}_{m]} - \mathcal{F}_{a}D^{a}H_{nkm} \\ \\ - \frac{1}{6}D_{[n}H^{abc}\mathcal{F}_{km]abc} - \frac{1}{2}H^{abc}D_{a}\mathcal{F}_{bcnkm} \\ \\ = 2(X^{a}D_{a}\mathcal{F}_{nkm} + D_{[n}X^{a}\mathcal{F}_{km]a}) + \beta^{G}_{a[n}\mathcal{F}^{a}_{\ \ km]} + \beta^{B}_{[nk}\mathcal{F}_{m]} - \frac{1}{2}\beta^{B}_{ab}\mathcal{F}^{ab}_{\ \ nkm} \end{array}$$

$$\begin{array}{l} \textbf{3)} \qquad \frac{D^{2}\mathcal{F}_{ijklm} - R_{a[i}\mathcal{F}^{a}{}_{jklm]} + R_{ab[ij}\mathcal{F}^{ab}{}_{klm]} + \frac{1}{4}(R - \frac{3}{4}H^{2})\mathcal{F}_{ijklm}}{+ \frac{1}{2}H^{abc}H_{ab[i}\mathcal{F}_{jklm]c} - \frac{1}{2}H^{abc}H_{a[ij}\mathcal{F}_{klm]bc}} \\ + \frac{1}{2}H^{abc}H_{a[ij}\mathcal{F}_{klm]} + H_{a[ij}D^{a}\mathcal{F}_{klm]} - \mathcal{F}_{a[ij}D^{a}H_{klm]} \\ + D^{a}H_{a[ij}\mathcal{F}_{klm]} + H_{a[ij}D^{a}\mathcal{F}_{klm]} - \mathcal{F}_{a[ij}D^{a}H_{klm]} \\ + \frac{1}{12}\varepsilon_{ijklmbdef}(D_{a}H^{abc}\mathcal{F}^{def} + H^{abc}D_{a}\mathcal{F}^{def} - \mathcal{F}^{abc}D_{a}H^{def}) = \\ = 2(X^{a}D_{a}\mathcal{F}_{ijklm} + D_{[i}X^{a}\mathcal{F}_{jklm]a}) + \beta^{G}_{a[i}\mathcal{F}^{a}{}_{jklm]} + \beta^{B}_{[ij}\mathcal{F}_{klm]} + \frac{1}{12}\varepsilon_{ijklmabcde}(\beta^{B})^{ab}\mathcal{F}^{cde} \end{array}$$

All three equations contain the expected Hodge-de Rham operator, vector X generates reparametrization and bosonic beta-functions appeared

T-duality as a origin of the *I*-deformation (NSNS case)

Analogously,
$$\bar{\beta}^X \equiv R - \frac{1}{12}H_{mnk}^2 + 4D^m X_m - 4X^m X_m = 0$$

The

With



and

future problems

☆ Does *I*-modified equations follow from kappa-symmetry. Yes!

[Wulff and Tseytlin '16]

What about other deformations, e.g. corresponding to the solutions of the CYBE. Yes!

[Yoshida et al. '16], [Hoare and van Tongeren '16]

 \checkmark What is the relation to the mirror model ?

[van Tongeren and G.A. '14], [Pachol and van Tongeren '16]

☆ Eta-deformed model is scale invariant but not Weyl invariant. Classical kappa-symmetry implies scale invariance only! Could this model still be used to define a critical string?

Eta-deformed model is related to the lambda-model by Poisson-Lie duality and analytic continuation. The latter model propagates in the IIB background and therefore is Weyl invariant. Extra fields? Non-locality?

[Yoshida et al. '17]

 \therefore Is there any way to obtain *I*-modified gravity equations from some Lagrangian?

 \Rightarrow What is *I*-modification for background fermionic fields?

/-modification destroys local supersymmetry. If there is still any local (hidden) symmetry?

 \checkmark What is a dual (non-commutative) gauge theory?

