

Conference on Symmetries, Geometry and Quantum Gravity
18-22 June 2018, Primošten, Croatia

Nonassociative differential geometry and gravity

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arXiv: 1710.11467

NC/NA geometry and gravity

Early Universe, singularities of BHs \Rightarrow QG \Rightarrow Quantum space-time
NC/NA space-time \Rightarrow Gravity on NC/NA spaces.

General Relativity (GR) is based on the **diffeomorphism symmetry**.
This concept (space-time symmetry) is difficult to generalize to
NC/NA spaces. Different approaches:

NC spectral geometry [Chamseddine, Connes, Marcolli '07; Chamseddine, Connes, Mukhanov '14].

Emergent gravity [Steinacker '10, '16].

Frame formalism, operator description [Burić, Madore '14; Fritz, Majid '16].

Twist approach [Wess et al. '05, '06; Ohl, Schenckel '09; Castellani, Aschieri '09; Aschieri, Schenkel '14].

NC gravity as a gauge theory of Lorentz/Poincaré group
[Chamseddine '01,'04, Cardela, Zanon '03, Aschieri, Castellani '09,'12; Dobrski '16].

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NA gravity: Motivation

Do **closed strings** provide a framework for **quantum gravity**?

In particular: closed strings propagating in a **locally non-geometric constant R -flux** background [Munich group '11].

The low energy limit: **Nonassociative gravity** on space-time.

Twist deformation: a well defined way to deform symmetries, in particular diffeomorphism symmetry, and the corresponding differential geometry [Wess et al. '06; Aschieri et al. '08,'09,...].

R -flux induced **cochain twist**: introduced in [Mylonas, Schupp, Szabo '14; Aschieri, Szabo '15].

NA gravity: General

NA gravity is based on:

- locally **non-geometric** constant R -flux.
- twist** \mathcal{F} and **associator** Φ with

$$\Phi(\mathcal{F} \otimes 1)(\Delta \otimes \text{id})\mathcal{F} = (1 \otimes \mathcal{F})(\text{id} \otimes \Delta)\mathcal{F} .$$

- equivariance (covariance) under the **twisted diffeomorphisms** (quasi-Hopf algebra of twisted diffeomorphisms).
- twisted differential geometry** in phase space. In particular: connection, curvature, torsion.
- NA Levi-Civita connection** and **projection** of phase space Einstein equations to space-time.

Our goals:

- consistently construct NA deformation of GR: NA Einstein equations, NA Einstein-Hilbert action; investigate phenomenological consequences.
- understand symmetries of the obtained NA gravity.

NA differential geometry: Review of twist deformation

Symmetry algebra g and the universal covering algebra Ug .

A well defined way of deforming symmetries: the **twist formalism**.

Twist \mathcal{F} (introduced by Drinfel'd in 1983-1985) is:

- an invertible element of $Ug \otimes Ug$
- fulfills the 2-cocycle condition (ensures the associativity of the \star -product).

$$\mathcal{F} \otimes 1(\Delta \otimes \text{id})\mathcal{F} = 1 \otimes \mathcal{F}(\text{id} \otimes \Delta)\mathcal{F}. \quad (2.1)$$

-additionally: $\mathcal{F} = 1 \otimes 1 + \mathcal{O}(\hbar)$; \hbar -deformation parameter.

Moyal-Weyl deformation: $\mathcal{F} = e^{-\frac{i}{2}\theta^{\mu\nu}\partial_\mu \otimes \partial_\nu}$

NC field theory [Szabo '03], NC Standard model [Wess et al. '01,...'06],

NC gravity [Chamseddine '01,'04; Cardela, Zanon '03; Wess et al. '05...].

Abelian twist: $\mathcal{F} = e^{-\frac{i}{2}\theta^{ab}X_a \otimes X_b}$,

where $X_a = X_a^\mu(x)\partial_\mu$, $[X^a, X^b] = 0$ and $\theta^{ab} = \text{const.}$. NC gravity [Aschieri, Castellani '09, ... '14].

NA differential geometry: R -flux induced cochain twist

Phase space \mathcal{M} : $x^A = (x^\mu, \tilde{x}_\mu = p_\mu)$, $\partial_A = (\partial_\mu, \tilde{\partial}^\mu = \frac{\partial}{\partial p_\mu})$.
2d dimensional, $A = 1, \dots, 2d$.

The twist \mathcal{F} :

$$\mathcal{F} = \exp\left(-\frac{i\hbar}{2}(\partial_\mu \otimes \tilde{\partial}^\mu - \tilde{\partial}^\mu \otimes \partial_\mu) - \frac{i\kappa}{2} R^{\mu\nu\rho} (p_\nu \partial_\rho \otimes \partial_\mu - \partial_\mu \otimes p_\nu \partial_\rho)\right), \quad (2.2)$$

with $R^{\mu\nu\rho}$ totally antisymmetric and constant, $\kappa := \frac{\ell_s^3}{6\hbar}$.

Does not fulfill the 2-cocycle condition

$$\Phi(\mathcal{F} \otimes 1)(\Delta \otimes \text{id})\mathcal{F} = (1 \otimes \mathcal{F})(\text{id} \otimes \Delta)\mathcal{F}. \quad (2.3)$$

The associator Φ :

$$\Phi = \exp(\hbar\kappa R^{\mu\nu\rho} \partial_\mu \otimes \partial_\nu \otimes \partial_\rho) =: \phi_1 \otimes \phi_2 \otimes \phi_3 = 1 \otimes 1 \otimes 1 + O(\hbar\kappa). \quad (2.4)$$

Notation: $\mathcal{F} = f^\alpha \otimes f_\alpha$, $\mathcal{F}^{-1} = \bar{f}^\alpha \otimes \bar{f}_\alpha$, $\Phi^{-1} =: \bar{\phi}_1 \otimes \bar{\phi}_2 \otimes \bar{\phi}_3$,

$$\mathcal{R} = \mathcal{F}^{-2} =: R^\alpha \otimes R_\alpha, \quad \mathcal{R}^{-1} = \mathcal{F}^2 =: \bar{R}^\alpha \otimes \bar{R}_\alpha.$$

Hopf algebra of diffeomorphisms $U\text{Vec}(\mathcal{M})$:

$$[u, v] = (u^B \partial_B v^A - v^B \partial_B u^A) \partial_A,$$

$$\Delta(u) = 1 \otimes u + u \otimes 1,$$

$$\epsilon(u) = 0, S(u) = -u.$$

Quasi-Hopf algebra of infinitesimal diffeomorphisms $U\text{Vec}^{\mathcal{F}}(\mathcal{M})$:

-algebra structure does not change

-coproduct is deformed: $\Delta^{\mathcal{F}} \xi = \mathcal{F} \Delta \mathcal{F}^{-1}$

-counit and antipod do not change: $\epsilon^{\mathcal{F}} = \epsilon, S^{\mathcal{F}} = S$.

On basis vectors:

$$\Delta_{\mathcal{F}}(\partial_{\mu}) = 1 \otimes \partial_{\mu} + \partial_{\mu} \otimes 1,$$

$$\Delta_{\mathcal{F}}(\tilde{\partial}^{\mu}) = 1 \otimes \tilde{\partial}^{\mu} + \tilde{\partial}^{\mu} \otimes 1 + i \kappa R^{\mu\nu\rho} \partial_{\nu} \otimes \partial_{\rho}.$$

NA differential geometry: NA tensor calculus

Guiding principle: Differential geometry on \mathcal{M} is covariant under $U\text{Vec}(\mathcal{M})$.

NA differential geometry on \mathcal{M} should be **covariant** under $U\text{Vec}^{\mathcal{F}}(\mathcal{M})$.

In practice: $U\text{Vec}(\mathcal{M})$ -module algebra \mathcal{A} (functions, forms, tensors) and $a, b \in \mathcal{A}$, $u \in \text{Vec}(\mathcal{M})$

$$u(ab) = u(a)b + au(b), \quad \text{Lie derivative, coproduct.}$$

The twist: $U\text{Vec}(\mathcal{M}) \rightarrow U\text{Vec}^{\mathcal{F}}(\mathcal{M})$ and $\mathcal{A} \rightarrow \mathcal{A}_\star$ with $ab \rightarrow a \star b = \bar{f}^\alpha(a) \cdot \bar{f}_\alpha(b)$. Then \mathcal{A}_\star is a $U\text{Vec}^{\mathcal{F}}(\mathcal{M})$ -module algebra:

$$\xi(a \star b) = \xi_{(1)}(a) \star \xi_{(2)}(b),$$

for $\xi \in U\text{Vec}^{\mathcal{F}}(\mathcal{M})$ and using the twisted coproduct $\Delta^{\mathcal{F}}\xi$.

Commutativity: $a \star b = \bar{f}^\alpha(a) \cdot \bar{f}_\alpha(b) = \bar{R}^\alpha(b) \star \bar{R}_\alpha(a) =: {}^\alpha b \star_\alpha a$

Associativity: $(a \star b) \star c = \phi_1 a \star (\phi_2 b \star \phi_3 c)$.

Functions: $C^\infty(\mathcal{M}) \rightarrow C^\infty(\mathcal{M})_\star$

$$\begin{aligned} f \star g &= \bar{f}^\alpha(f) \cdot \bar{f}_\alpha(g) & (2.5) \\ &= f \cdot g + \frac{i\hbar}{2} (\partial_\mu f \cdot \tilde{\partial}^\mu g - \tilde{\partial}^\mu f \cdot \partial_\mu g) + i\kappa R^{\mu\nu\rho} p_\nu \partial_\rho f \cdot \partial_\mu g + \dots, \end{aligned}$$

$$[x^\mu \star, x^\nu] = 2i\kappa R^{\mu\nu\rho} p_\rho, \quad [x^\mu \star, p_\nu] = i\hbar \delta^\mu_\nu, \quad [p_\mu \star, p_\nu] = 0,$$

$$[x^\mu \star, x^\nu \star, x^\rho] = \ell_s^3 R^{\mu\nu\rho}.$$

Forms: $\Omega^\sharp(\mathcal{M}) \rightarrow \Omega^\sharp(\mathcal{M})_\star$

$$\omega \wedge_\star \eta = \bar{f}^\alpha(\omega) \wedge \bar{f}_\alpha(\eta), \quad (2.6)$$

$$f \star dx^A = dx^C \star (\delta^A_C f - i\kappa \mathcal{R}^{AB}{}_C \partial_B f),$$

with non-vanishing components $\mathcal{R}^{x^\mu, x^\nu}{}_{\tilde{x}^\rho} = R^{\mu\nu\rho}$. Basis 1-forms

$$\begin{aligned} (dx^A \wedge_\star dx^B) \wedge_\star dx^C &= \phi_1(dx^A) \wedge_\star (\phi_2(dx^B) \wedge_\star \phi_3(dx^C)) \\ &= dx^A \wedge_\star (dx^B \wedge_\star dx^C) = dx^A \wedge dx^B \wedge dx^C. \end{aligned}$$

NA tensor calculus: duality, derivation, Lie derivative

Exterior derivative d : $d^2 = 0$ and the undeformed Leibniz rule

$$d(\omega \wedge_{\star} \eta) = d\omega \wedge_{\star} \eta + (-1)^{|\omega|} \omega \wedge_{\star} d\eta. \quad (2.7)$$

Duality, \star -pairing:

$$\langle \omega, u \rangle_{\star} = \langle \bar{f}^{\alpha}(\omega), \bar{f}_{\alpha}(u) \rangle. \quad (2.8)$$

\star Lie derivative:

$$\mathcal{L}_u^{\star}(T) = \mathcal{L}_{\bar{f}_{\alpha}(u)}(\bar{f}^{\alpha}(T)), \quad (2.9)$$

$$[\mathcal{L}_u^{\star}, \mathcal{L}_v^{\star}]_{\bullet} = \mathcal{L}_{[u, v]_{\star}},$$

$$\mathcal{L}_u^{\star}(\omega \wedge_{\star} \eta) = \mathcal{L}_{\bar{f}_{\alpha}(u)}^{\star}(\bar{\phi}_2^{\alpha} \omega) \wedge_{\star} \bar{\phi}_3 \eta + \alpha(\bar{\phi}_1 \bar{\phi}_2 u) \wedge_{\star} \mathcal{L}_{\alpha(\bar{\phi}_2 \bar{\phi}_3 u)}^{\star}(\bar{\phi}_3 \bar{\phi}_3 \eta),$$

with $[u, v]_{\star} = [\bar{f}^{\alpha}(u), \bar{f}_{\alpha}(v)]$. Relation of \mathcal{L}_u^{\star} with diffeomorphism symmetry in space-time needs to be understood.

NA differential geometry: NA connection

\star -connection:

$$\begin{aligned}\nabla^\star &: \text{Vec}_\star &\longrightarrow & \text{Vec}_\star \otimes_\star \Omega_\star^1 \\ u &\longmapsto & \nabla^\star u, & \end{aligned} \quad (2.10)$$

$$\nabla^\star(u \star f) = (\bar{\phi}_1 \nabla^\star(\bar{\phi}_2 u)) \star \bar{\phi}_3 f + u \otimes_\star df, \quad (2.11)$$

the right Leibniz rule, for $u \in \text{Vec}_\star$ and $f \in A_\star$. In particular:

$$\begin{aligned}\nabla^\star \partial_A &=: \partial_B \otimes_\star \Gamma_A^B =: \partial_B \otimes_\star (\Gamma_{AC}^B \star dx^C). & (2.12) \\ d_{\nabla^\star}(\partial_A \otimes_\star \omega^A) &= \partial_A \otimes_\star (d\omega^A + \Gamma_B^A \wedge_\star \omega^B),\end{aligned}$$

for $\omega^A \in \Omega_\star^\sharp$.

Connection on forms, dual connection $\star \nabla \omega$:

$$\begin{aligned}\star \nabla &: \Omega_\star^1 &\longrightarrow & \Omega_\star^1 \otimes_\star \Omega_\star^1, \\ \omega &\longmapsto & \star \nabla \omega, \\ \star \nabla(f \star \omega) &= & \phi_1 f \star (\phi_3 \star \nabla(\phi_2 \omega)) + df \otimes_\star \omega. & (2.13)\end{aligned}$$

the left Leibniz rule, for $\omega \in \Omega_\star^1$ and $f \in A_\star$. In particular:

$$d_{\star \nabla}(\omega_A \otimes_\star dx^A) = (d\omega_A - \omega_B \wedge_\star \Gamma_A^B) \otimes_\star dx^A.$$

NA differential geometry: NA torsion, NA curvature

Torsion:

$$\begin{aligned} T^* &:= d_{\nabla^*} (\partial_A \otimes_{\star} dx^A) : \text{Vec}_{\star} \otimes_{\star} \text{Vec}_{\star} \rightarrow \text{Vec}_{\star}, \\ T^*(\partial_A, \partial_B) &= \partial_C \star (\Gamma_{AB}^C - \Gamma_{BA}^C) =: \partial_C \star T_{AB}^C. \end{aligned}$$

Torsion-free condition: $\Gamma_{AB}^C = \Gamma_{BA}^C$.

Curvature:

$$\begin{aligned} R^* &:= d_{\nabla^*} \bullet d_{\nabla^*} : \text{Vec}_{\star} \longrightarrow \text{Vec}_{\star} \otimes_{\star} \Omega_{\star}^2, \\ R^*(\partial_A) &= \partial_C \otimes_{\star} (d\Gamma_A^C + \Gamma_B^C \wedge_{\star} \Gamma_A^B) = \partial_C \otimes_{\star} R_A^C, \end{aligned}$$

with the curvature coefficients

$$\begin{aligned} R^*(\partial_A, \partial_B, \partial_C) &= \langle \partial_D \otimes_{\star} R_A^D, \partial_B \wedge_{\star} \partial_C \rangle_{\star} \\ &= \partial_D \star (\partial_C \Gamma_{AB}^D - \partial_B \Gamma_{AC}^D - \Gamma_{B'E}^D \star (\delta^E_B \Gamma_{AC}^{B'} + i\kappa \mathcal{R}^{EG}_B (\partial_G \Gamma_{AC}^{B'})) \\ &\quad + \Gamma_{B'E}^D \star (\delta^E_C \Gamma_{AB}^{B'} + i\kappa \mathcal{R}^{EG}_C (\partial_G \Gamma_{AB}^{B'}))) \\ &= \partial_D \star R^D_{ABC}. \end{aligned} \tag{2.14}$$

First Cartan structure equation:

$$T^*(u, v) = \phi_1 \nabla_{\phi_2 v}^* (\phi_3 u) - \phi_1 \nabla_{\phi_2 \alpha u}^* (\phi_3 \alpha v) + [u, v]_{\star}. \quad (2.15)$$

Second Cartan structure equation:

$$\begin{aligned} R^*(z, u, v) = & \kappa_1 \check{\phi}_1 \phi_1' \nabla_{\bar{\rho}_3 \bar{\zeta}_3 \bar{\phi}_3 \phi_3' v}^* (\bar{\rho}_1 \bar{\phi}_1 \kappa_2 \check{\phi}_2 \phi_2' \nabla_{\bar{\rho}_2 \bar{\zeta}_2 \check{\phi}_3 u}^* \bar{\zeta}_1 \bar{\phi}_2 \kappa_3 z) \\ & - \kappa_1 \check{\phi}_1 \phi_1' \nabla_{\bar{\rho}_3 \bar{\zeta}_3 \bar{\phi}_3 \phi_3' \alpha u}^* (\bar{\rho}_1 \bar{\phi}_1 \kappa_2 \check{\phi}_2 \phi_2' \nabla_{\bar{\rho}_2 \bar{\zeta}_2 \check{\phi}_3 \alpha v}^* \bar{\zeta}_1 \bar{\phi}_2 \kappa_3 z) + \nabla_{[u, v]_{\star}}^* z. \end{aligned}$$

Bianchi identities:

$$\begin{aligned} dT^A + \Gamma_B^A \wedge_{\star} T^B &= R_B^A \wedge_{\star} dx^B, \quad (2.16) \\ dR_A^C + \Gamma_B^C \wedge_{\star} R_A^B - R_B^C \wedge_{\star} \Gamma_A^B &= \\ &= \Gamma_B^C \wedge_{\star} (\Gamma_D^B \wedge_{\star} \Gamma_A^D) - \phi_1 \Gamma_B^C \wedge_{\star} (\phi_2 \Gamma_D^B \wedge_{\star} \phi_3 \Gamma_A^D). \end{aligned}$$

Ricci tensor:

$$\begin{aligned} \text{Ric}^*(u, v) &:= -\langle R^*(u, v, \partial_A), dx^A \rangle_* \\ \text{Ric}^* &= \text{Ric}_{AD} \star (dx^D \otimes_* dx^A). \end{aligned} \quad (2.17)$$

Components from $\text{Ric}_{BC} := \text{Ric}^*(\partial_B, \partial_C)$

$$\begin{aligned} \text{Ric}_{BC} &= \partial_A \Gamma_{BC}^A - \partial_C \Gamma_{BA}^A + \Gamma_{B'A}^A \star \Gamma_{BC}^{B'} - \Gamma_{B'C}^A \star \Gamma_{BA}^{B'} \\ &+ i\kappa \Gamma_{B'E}^A \star (\mathcal{R}^{EG}_A (\partial_G \Gamma_{BC}^{B'}) - \mathcal{R}^{EG}_C (\partial_G \Gamma_{BA}^{B'})) \\ &+ i\kappa \mathcal{R}^{EG}_A \partial_G \partial_C \Gamma_{BE}^A - i\kappa \mathcal{R}^{EG}_A \partial_G (\Gamma_{B'E}^A \star \Gamma_{BC}^{B'} - \Gamma_{B'C}^A \star \Gamma_{BE}^{B'}) \\ &+ \kappa^2 \mathcal{R}^{AF}_D (\mathcal{R}^{EG}_A \partial_F (\Gamma_{B'E}^D \star \partial_G \Gamma_{BC}^{B'}) - \mathcal{R}^{EG}_C \partial_F (\Gamma_{B'E}^D \star \partial_G \Gamma_{BA}^{B'})) . \end{aligned} \quad (2.18)$$

Scalar curvature cannot be defined along these lines: cannot be seen as a map and inverse metric tensor needed. Not straightforward.

NA deformation of GR: NA Levi-Civita connection

GR connection $\Gamma_{\mu\nu}^{\text{LC}\rho}$ is a **Levi-Civita** connection: torsion-free and metric compatible $\nabla_{\alpha}g_{\mu\nu} = 0$. Generalization to our NA phase space:

Metric tensor $g^{\star} \in \Omega_{\star}^1 \otimes_{\star} \Omega_{\star}^1$ and $g^{\star}(u, v) = \langle g^{\star}, u \otimes_{\star} v \rangle_{\star} \in C^{\infty}(\mathcal{M})_{\star}$. In the coordinate basis:

$$g^{\star} = g_{AB} \star (dx^A \otimes_{\star} dx^B), \quad g^{\star}(\partial_A, \partial_B) = g_{AB} = g_{BA}.$$

Metric compatibility condition: $\star\nabla g^{\star} = 0$. We calculate

$$\begin{aligned} dg_{AB} &= d\langle g^{\star}, \partial_A \otimes_{\star} \partial_B \rangle_{\star} \\ &= \langle \star\nabla g^{\star}, \partial_A \otimes_{\star} \partial_B \rangle_{\star} + \langle \phi_1 g^{\star}, \phi_2 \nabla^{\star}(\phi_3(\partial_A \otimes_{\star} \partial_B)) \rangle_{\star} \\ &= \langle g^{\star}, \nabla^{\star}(\partial_A \otimes_{\star} \partial_B) \rangle_{\star}. \end{aligned}$$

Calculation straightforward up to

$$G_{CN} \star \Gamma_{AD}^N = \frac{1}{2} (\partial_D g_{AC} + \partial_A g_{DC} - \partial_C g_{AD} + i\kappa \mathcal{R}^{EF}{}_C (\partial_E \partial_D g_{AF} + \partial_E \partial_A g_{DF})), \quad (3.19)$$

$G_{MN} = g_{MN} + i\kappa \mathcal{R}^{EF}{}_M \partial_E g_{NF}$. Problems with straightforward inversion noted already in [Blumenhagen, Fuchs '16]. Working with $G^{MC} \star G_{CN} = \delta_N^M$ leads to:

$$G^{MC} \star (G_{CN} \star \Gamma_{AD}^N) = (\bar{\phi}_1 G^{MC} \star \bar{\phi}_2 G_{CN}) \star \bar{\phi}_3 \Gamma_{AD}^N \neq \Gamma_{AD}^M. \quad (3.20)$$

Inversion **possible** in terms of differential operators:

$$\Gamma_{AD}^S = G^{*SC} * W_{CAD} + \sum_{\vec{\lambda}} \frac{(i\kappa)^{|\vec{\lambda}|}}{\vec{\lambda}!} (-1)^{l(\vec{\lambda})} Y_G^{(\vec{\lambda})S}{}_M (G^{*MC} * W_{CAD}),$$

with $*$ being the \star -product corresponding to the twist

$$F = \exp \left(-\frac{i\hbar}{2} (\partial_\mu \otimes \tilde{\partial}^\mu - \tilde{\partial}^\mu \otimes \partial_\mu) \right)$$

and $G^{*SC} * G_{CN} = \delta_N^S$.

Also:

$$W_{CAD} = \frac{1}{2} (\partial_D g_{AC} + \partial_A g_{DC} - \partial_C g_{AD} + i\kappa \mathcal{R}^{EF}{}_C (\partial_E \partial_D g_{AF} + \partial_E \partial_A g_{DF})),$$

$$Y_G{}^M{}_N = \bar{f}_R^\beta (G^{*MC} * \bar{f}_R^\alpha (G_{CN})) \bar{f}_{R\beta} \bar{f}_{R\alpha} =: \sum_{n=0}^{\infty} \frac{(i\kappa)^n}{n!} Y_G^{(n)M}{}_N,$$

$$Y_G^{(0)M}{}_N = \delta_N^M,$$

$$Y_G^{(1)M}{}_N = \frac{1}{2} \bar{f}_R^\beta (G^{*MC} * (R^{\mu\nu\rho} p_\nu \partial_\rho G_{CN})) \bar{f}_{R\beta} \partial_\mu - \frac{1}{2} \bar{f}_R^\beta (G^{*MC} * (\partial_\mu G_{CN})) \bar{f}_{R\beta} R^{\mu\nu\rho} p_\nu \partial_\rho,$$

$$\text{with } F_R^{-1} = \exp\left(\frac{i\kappa}{2} R^{\mu\nu\rho} (p_\nu \partial_\rho \otimes \partial_\mu - \partial_\mu \otimes p_\nu \partial_\rho)\right) = \bar{f}_R^\alpha \otimes \bar{f}_{R\alpha}.$$

Connection coefficients, expanded up to first order in $\hbar\kappa$:

$$\Gamma_{AD}^{S(0,0)} = \Gamma_{AD}^{LCS} = \frac{1}{2} g^{SQ} (\partial_D g_{AQ} + \partial_A g_{DQ} - \partial_Q g_{AD}), \quad (3.21)$$

$$\Gamma_{AD}^{S(0,1)} = -\frac{i\hbar}{2} g^{SP} ((\partial_\mu g_{PQ}) \tilde{\partial}^\mu \Gamma_{AD}^{LCQ} - (\tilde{\partial}^\mu g_{PQ}) \partial_\mu \Gamma_{AD}^{LCQ}),$$

$$\Gamma_{AD}^{S(1,0)} = i\kappa R^{\alpha\beta\gamma} \left(\tilde{g}^S_\gamma g_{\beta N} (\partial_\alpha \Gamma_{AD}^{LCN}) - g^{SM} p_\beta (\partial_\gamma g_{MN}) \partial_\alpha \Gamma_{AD}^{LCN} \right),$$

$$\Gamma_{AD}^{S(1,1)} = \frac{\hbar\kappa}{2} R^{\alpha\beta\gamma} \left[\dots \text{long expression} \dots \right. \\ \left. + (\partial_\alpha g^{SQ}) (\partial_\beta g_{QP}) \partial_\gamma \Gamma_{AD}^{LCP} \right].$$

Comments:

$-\Gamma_{AD}^{S(0,1)}$ and $\Gamma_{AD}^{S(1,0)}$ imaginary, $\Gamma_{AD}^{S(1,1)}$ real.

-for g_{MN} that does not depend on the momenta p_μ , only the last term in $\Gamma_{AD}^{S(1,1)}$ remains.

$-\tilde{g}^S_\gamma = g^{SM} \delta_{M, \tilde{x}_\gamma}$.

NA deformation of GR: NA vacuum Einstein equation

We can write **vacuum Einstein equations** in phase space as:

$$\text{Ric}_{BC} = 0 . \quad (3.22)$$

Our strategy: expand Ricci tensor (2.17) in term of (3.21), i. e. the metric tensor g_{MN} . This gives **Einstein equations in phase space**. How do we obtain the induced equations in space-time?

From phase space to space-time:

- ▶ start from objects in space-time M $g = g_{\mu\nu} dx^\mu \otimes dx^\nu$ and lift them to phase space \mathcal{M} foliated with **leaves of constant momenta**, each leaf is diffeomorphic to M .

$$\begin{array}{ccc}
 C^\infty(\mathcal{M}) & \xrightarrow{Q} & \widehat{C^\infty(\mathcal{M})} \\
 \uparrow \pi^* & & \downarrow s_{\bar{p}}^* = \sigma^* \\
 C^\infty(M) & \xrightarrow{Q_{\bar{p}}} & \widehat{C^\infty(M)}
 \end{array}$$

Functions, forms: pullback using the canonical projection $\pi : \mathcal{M} \rightarrow M$.

Vector fields: $v^\mu(x) \partial_\mu \mapsto \pi^*(v^\mu)(x, p) \partial_\mu$ and $\pi^*(v^\mu)(x, p) = (v^\mu)(\pi(x, p)) = v^\mu(x)$.

Metric tensor: $g = g_{\mu\nu} dx^\mu \otimes dx^\nu \rightarrow \hat{g}_{MN} dx^M \otimes dx^N$ with

$$(\hat{g}_{MN}(x)) = \begin{pmatrix} g_{\mu\nu}(x) & 0 \\ 0 & h^{\mu\nu}(x) \end{pmatrix}. \quad (3.23)$$

Note the additional nondegenerate bilinear $h(x)^{\mu\nu} d\tilde{x}_\mu \otimes d\tilde{x}_\nu$;
natural choice $h(x)^{\mu\nu} = \eta^{\mu\nu}$.

- ▶ Do all calculations in phase space, using the twisted differential geometry. In particular, calculate Ric_{BC} in terms of g_{AB} , (2.17), (3.21).
- ▶ Finally, project the result to space-time using the zero section $x \mapsto \sigma(x) = (x, 0)$.

Functions, forms: pullback to the zero momentum leaf:

Vector fields: $v^\mu(x, p) \partial_\mu + \tilde{v}_\mu(x, p) \tilde{\partial}^\mu \mapsto v^\mu(x, 0) \partial_\mu$.

Ricci tensor: $\text{Ric} \rightarrow \text{Ric}^{*\circ} = \text{Ric}_{\mu\nu}^\circ dx^\mu \otimes dx^\nu$,
 $\text{Ric}_{\mu\nu}^\circ(x) = \sigma^*(\text{Ric}_{\mu\nu})(x, p) = \text{Ric}_{\mu\nu}(x, 0)$.

NA deformation of GR: NA gravity in space-time

The lifted metric $\hat{g}_{MN} dx^M \otimes dx^N = g_{MN} \star (dx^M \otimes_{\star} dx^N)$,

$$g_{MN}(x) = \begin{pmatrix} g_{\mu\nu}(x) & \frac{i\kappa}{2} R^{\sigma\nu\alpha} \partial_{\sigma} g_{\mu\alpha} \\ \frac{i\kappa}{2} R^{\sigma\mu\alpha} \partial_{\sigma} g_{\alpha\nu} & \eta^{\mu\nu}(x) \end{pmatrix}. \quad (3.24)$$

Ricci tensor in space-time, (expanded up to first order in $\hbar\kappa$):

$$\begin{aligned} \text{Ric}^{\circ}_{\mu\nu} = & \text{Ric}^{\text{LC}}_{\mu\nu} + \frac{\ell_s^3}{12} R^{\alpha\beta\gamma} \left(\partial_{\rho} (\partial_{\alpha} g^{\rho\sigma} (\partial_{\beta} g_{\sigma\tau}) \partial_{\gamma} \Gamma^{\text{LC}}_{\mu\nu}{}^{\tau}) \right. \\ & \left. - \partial_{\nu} (\partial_{\alpha} g^{\rho\sigma} (\partial_{\beta} g_{\sigma\tau}) \partial_{\gamma} \Gamma^{\text{LC}}_{\mu\rho}{}^{\tau}) \right. \\ & + \partial_{\gamma} g_{\tau\omega} (\partial_{\alpha} (g^{\sigma\tau} \Gamma^{\text{LC}}_{\sigma\nu}{}^{\rho}) \partial_{\beta} \Gamma^{\text{LC}}_{\mu\rho}{}^{\omega} - \partial_{\alpha} (g^{\sigma\tau} \Gamma^{\text{LC}}_{\sigma\rho}{}^{\omega}) \partial_{\beta} \Gamma^{\text{LC}}_{\mu\nu}{}^{\omega} \\ & + (\Gamma^{\text{LC}}_{\mu\rho}{}^{\sigma} \partial_{\alpha} g^{\rho\tau} - \partial_{\alpha} \Gamma^{\text{LC}}_{\mu\rho}{}^{\sigma} g^{\rho\tau}) \partial_{\beta} \Gamma^{\text{LC}}_{\sigma\nu}{}^{\omega} \\ & \left. - (\Gamma^{\text{LC}}_{\mu\nu}{}^{\sigma} \partial_{\alpha} g^{\rho\tau} - \partial_{\alpha} \Gamma^{\text{LC}}_{\mu\nu}{}^{\sigma} g^{\rho\tau}) \partial_{\beta} \Gamma^{\text{LC}}_{\sigma\rho}{}^{\omega} \right). \end{aligned} \quad (3.25)$$

Vacuum Einstein equations in space-time:

$$\text{Ric}^{\circ}_{\mu\nu} = 0. \quad (3.26)$$

NA deformation of GR: Comments

- ▶ R -flux (via NA differential geometry) generates non-trivial dynamical consequences on spacetime, they are independent of \hbar and real-valued.
- ▶ the linear R -flux correction to $\text{Ric}^\circ_{\mu\nu}$ is not a total derivative.
- ▶ Why zero momentum leaf? Pulling back to a leaf of constant momentum $p = p^\circ$ (generally) gives a non-vanishing imaginary contribution $\text{Ric}^{(1,0)}_{\mu\nu} \Big|_{p=p^\circ}$ to the spacetime Ricci tensor. Also, n -triproducts calculated on the zero momentum leaf [Aschieri, Szabo '15] coincide with those proposed in [Munich group '11].
- ▶ Why $h(x)^{\mu\nu} = \eta^{\mu\nu}$? The simplest choice, can be extended. In relation with Born geometry [Freidel et al. '14]: in our model nonassociativity does not generate curved momentum space. Investigate $h(x)^{\mu\nu} \neq \eta^{\mu\nu} \dots$

Discussion

Our goals:

- ▶ Phenomenological consequences (R -flux induced corrections to GR solutions): to be investigated.
- ▶ Construction of scalar curvature, matter fields, full Einstein equations: to be investigated.
- ▶ Twisted diffeomorphism symmetry and "quantum diffeomorphisms": to be understood better.
- ▶ NA gravity as a gauge theory of Lorentz symmetry, NA Einstein-Cartan gravity: better understanding of NA gauge symmetry is needed, L_∞ algebra? See also [Blumenhagen et al. '18].